# Relativistic thermodynamics and magneto-hydrodynamics 

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## Motivation

Victoria

Hydrodynamics


## Motivation

Flow of hot sub-nuclear matter: heavy-ion collisions

Flow of dense subnuclear matter: neutron star mergers

First few microseconds after the Big Bang

I didn't understand hydrodynamics as a student

## Motivation

THE SUN'S ATMOSPHERE ISA SUPERHOT PLASMA GOVERNED BY MAGNETOHYDRODYNAMIC FORCES...


AH, YES, OF COURSE.


WHENEVER I HEAR THE WORD
"MAGNETOHYDRODYNAMIC" MY BRAIN JUST REPLACES IT WITH "MAGIC."

# To understand hydrodynamics, first understand thermodynamics 

## Thermodynamics

System in external time-independent $g_{\mu v}, A_{\mu}$

Compute $W=-i \ln Z\left[g_{\mu \nu}, A_{\mu}\right]$

Local correlations $\Rightarrow W[g, A]=\int d^{d+1} x \sqrt{-g} \mathcal{F}(g, A)$

Near-uniform fields $\Rightarrow$ expand $\mathcal{F}(g, A)$ in derivatives of $\mathrm{g}, \mathrm{A}$

Leading order $\Rightarrow \mathcal{F}(g, A)=P+O(\partial)$

## Thermodynamic variables

Timelike Killing vector $\mathrm{V}^{\mu}$, e.g. $\mathrm{V}^{\mu}=(1, \mathbf{0})$ for matter "at rest"

$$
T=\frac{1}{\beta_{0} \sqrt{-V^{2}}}, \quad u^{\mu}=\frac{V^{\mu}}{\sqrt{-V^{2}}}, \quad \mu=\frac{V^{\mu} A_{\mu}+\Lambda_{V}}{\sqrt{-V^{2}}}
$$

JLY arXiv:1310.7024

Definition of electric and magnetic fields:

$$
F_{\mu \nu}=u_{\mu} E_{\nu}-u_{\nu} E_{\mu}-\epsilon_{\mu \nu \rho \sigma} u^{\rho} B^{\sigma}
$$

## Equilibrium relations

$$
u^{\lambda} \partial_{\lambda} T=0, \quad u^{\lambda} \partial_{\lambda} \mu=0
$$

$$
a_{\lambda}=-\partial_{\lambda} T / T
$$

$$
E^{\alpha}-T \Delta^{\alpha \beta} \partial_{\beta}\left(\frac{\mu}{T}\right)=0
$$

$$
\nabla_{\mu} u_{\nu}=-u_{\mu} a_{\nu}-\frac{1}{2} \epsilon_{\mu \nu \alpha \beta} u^{\alpha} \Omega^{\beta}
$$

things don't depend on time
gravitational potential induces temperature gradient
electric field induces charge gradient: this is electric screening
shear and expansion vanish in equilibrium

$$
\begin{aligned}
a^{\mu} & \equiv u^{\lambda} \nabla_{\lambda} u^{\mu} \\
\Omega^{\mu} & \equiv \epsilon^{\mu \nu \alpha \beta} u_{\nu} \nabla_{\alpha} u_{\beta}
\end{aligned}
$$

## Bound charges and bound currents

$$
\delta_{A, F} W=\int d^{d+1} x \sqrt{-g}\left[J_{\mathrm{f}}^{\mu} \delta A_{\mu}+\frac{1}{2} M^{\mu \nu} \delta F_{\mu \nu}\right]
$$

The separation of $J_{f}$ and $M$ is ambiguous.
But the total current is not:

$$
J^{\mu}=J_{\mathrm{f}}^{\mu}-\nabla_{\lambda} M^{\lambda \mu}
$$

"free current" "bound current"

Can fix the ambiguity by trading $\partial_{\alpha} \mu$ for $E_{\alpha}$.
Then $J_{\mathrm{f}}^{\mu}=\rho u^{\mu} \quad$ where $\quad \rho \equiv \partial \mathcal{F} / \partial \mu$

## Bound charges and bound currents

Define charge density and spatial current:

Polarization vectors:

$$
M_{\mu \nu}=p_{\mu} u_{\nu}-p_{\nu} u_{\mu}-\epsilon_{\mu \nu \rho \sigma} u^{\rho} m^{\sigma}
$$

$$
\begin{aligned}
& \mathcal{N}=\rho-\nabla_{\mu} p^{\mu}+p^{\mu} a_{\mu}-m_{\mu} \Omega^{\mu} \\
& \mathcal{J}^{\mu}=\epsilon^{\mu \nu \rho \sigma} u_{\nu} \nabla_{\rho} m_{\sigma}+\epsilon^{\mu \nu \rho \sigma} u_{\nu} a_{\rho} m_{\sigma}
\end{aligned}
$$

$\mathrm{a}_{\mu}=$ acceleration $\Omega_{\mu}=$ vorticity

## Bound charges and bound currents

Define charge density and spatial current:

Polarization vectors:

$$
M_{\mu \nu}=p_{\mu} u_{\nu}-p_{\nu} u_{\mu}-\epsilon_{\mu \nu \rho \sigma} u^{\rho} m^{\sigma}
$$

$$
\begin{aligned}
& n=\rho-\boldsymbol{\nabla} \cdot \mathbf{p}-\mathbf{p} \cdot \boldsymbol{\nabla} T / T-2 \mathbf{m} \cdot \boldsymbol{\omega} \\
& \mathbf{J}=\boldsymbol{\nabla} \times \mathbf{m}+\mathbf{m} \times \boldsymbol{\nabla} T / T
\end{aligned}
$$

These were equilibrium charges and currents.
Now need to find equilibrium $\mathrm{T}^{\mu \nu}$.
For that, need the derivative expansion.

## Derivative expansion

$$
W[g, A]=\int \sqrt{-g} p+O(\partial)
$$

How do we count derivatives?
Clearly, $g_{\mu v}, T \sim O(1)$
In equilibrium, $E^{\alpha}-T \Delta^{\alpha \beta} \partial_{\beta}\left(\frac{\mu}{T}\right)=0$
So if $\mu \sim \mathrm{O}(1)$, then $\mathrm{E} \sim \mathrm{O}(\partial)$. This is screening.
No similar constraint on B , can take $\mathrm{B} \sim \mathrm{O}(\partial)$ or $\mathrm{B} \sim \mathrm{O}(1)$

## Derivative expansion

$$
W[g, A]=\int \sqrt{-g} p+O(\partial)
$$

Weak E, B: $p=p(T, \mu)$
Insulator in strong $E, B$ fields: $p=p\left(T, E^{2}, B^{2}, E \cdot B\right)$
Conductor in strong B-field: $\mathrm{p}=\mathrm{p}\left(\mathrm{T}, \mu, \mathrm{B}^{2}\right)$

## Example: P-invariant conductor in strong B field

Free energy: $\quad \mathcal{F}(g, A)=p\left(T, \mu, B^{2}\right)+M_{\Omega}\left(T, \mu, B^{2}\right) B \cdot \Omega+O\left(\partial^{2}\right)$

Vary $W[g, A]=\int d^{d+1} x \sqrt{-g} \mathcal{F}(g, A)$ to find $T^{\mu v}, J \mu$

In constant B-field: $\quad T_{s}^{\mu \nu}=Q_{s}^{\mu} u^{\nu}+Q_{s}^{\nu} u^{\mu}, \quad Q_{s}^{\alpha}=M_{\Omega} \epsilon^{\alpha \mu \nu \rho} u_{\mu} B_{\nu} n_{\rho}$


Angular momentum:

$$
\frac{\mathbf{L}}{V}=2 M_{\Omega} \mathbf{B}
$$

## Example: P-invariant conductor in strong B field

System at rest
in flat space, constant B-field:

$$
\frac{\mathbf{L}}{V}=2 M_{\Omega} \mathbf{B}
$$

System rotating in flat space, $\mathbf{m}=2 M_{\Omega} \boldsymbol{\omega}$ no B-field:

## Fluid with a global $\mathrm{U}(1)$

$$
W[g, A]=\int d^{4} x \sqrt{-g}\left[p(T, \mu)+\sum_{n} f_{n}(T, \mu) s_{n}^{(2)}\right]+\ldots
$$

| $n$ | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $s_{n}^{(2)}$ | $R$ | $a^{2}$ | $\Omega^{2}$ | $B^{2}$ | $B \cdot \Omega$ | $E^{2}$ | $E \cdot a$ | $B \cdot E$ | $B \cdot a$ |
| P | + | + | + | + | + | + | + | - | - |
| C | + | + | + | + | - | + | - | + | - |
| T | + | + | + | + | + | + | + | - | - |
| W | n/a | n/a | 2 | 4 | 3 | 4 | n/a | 4 | n/a |

Nine thermodynamic susceptibilities $f_{n}(T, \mu)$, have to be computed from the microscopics, just like $p(T, \mu)$

## Fluid with a global $\mathrm{U}(1)$

$$
W[g, A]=\int d^{4} x \sqrt{-g}\left[p(T, \mu)+\sum_{n} f_{n}(T, \mu) s_{n}^{(2)}\right]+\ldots
$$

$\mathrm{f}_{1}: T$ - and $\mu$-dependent Newton's constant
$\mathrm{f}_{2}$ : pressure response to $(\nabla \mathrm{T})^{2}$
$f_{3}$ : pressure response to (vorticity) ${ }^{2}$
$f_{4,6,8}$ : magnetic, electric, and magneto-electric suseptibilities
$f_{5}$ : magneto-vortical susceptibility, determines $\mathbf{L} \sim \mathbf{B}, \mathbf{m} \sim \boldsymbol{\omega}$
$\mathrm{f}_{7,9}$ : pressure response to $\mathbf{E} \cdot \mathbf{\nabla} \mathbf{T}, \mathbf{B} \cdot \mathbf{\nabla} \mathbf{T}$

## Example: no external E,B fields

QCD with $\boldsymbol{\mu}_{\mathrm{B}} \neq \mathbf{0}$ : vary $\mathrm{W}[\mathrm{g}, \mathrm{A}]$, get $T^{\mu v}$ and $\mathrm{J}^{\mu}$ in terms of five susceptibilities $f_{n}(T, \mu), n=1,2,3,5,7$ besides the pressure $p(T, \mu)$

CFT with $\boldsymbol{\mu} \neq \mathbf{0}$ : vary $\mathrm{W}[\mathrm{g}, \mathrm{A}]$, get $\mathrm{T}^{\mu v}$ and $\mathrm{J} \mathrm{\mu}$ in terms of three susceptibilities $f_{n}(T, \mu), n=1,3,5$ besides the pressure $p(T, \mu)$

Various combinations of $f_{n}(T, \mu)$ and their derivatives in $T^{\mu v}, J \mu$ are often called "thermodynamic transport coefficients".

Can be computed perturbatively, on the lattice, or in AdS/CFT BRSSS 0712.2451, Romatschke, Son 0903.3946, Moore, Sohrabi 1007.5333, 1210.3340, Arnold, Vaman, Wu, Xiao 1105.4645, Philipsen, Schäfer 1311.6618, Megias, Valle 1408.0165, Finazzo, Rougemont, Marrochio, Noronha 1412.2968, Buzzegoli, Grossi, Becattini 1704.02808

## If you really want to see the expressions

$$
\begin{aligned}
T^{\mu \nu} & =\mathcal{E} u^{\mu} u^{\nu}+\mathcal{P} \Delta^{\mu \nu}+\mathcal{Q}^{\mu} u^{\nu}+\mathcal{Q}^{\nu} u^{\mu}+\mathcal{T}^{\mu \nu} \\
J^{\mu} & =\mathcal{N} u^{\mu}+\mathcal{J}^{\mu}
\end{aligned}
$$

$$
\mathcal{E}=\epsilon+\left(f_{1}^{\prime}-f_{1}\right) R+\left(4 f_{1}^{\prime}+2 f_{1}^{\prime \prime}-f_{2}-f_{2}^{\prime}\right) a^{2}
$$

$$
+\left(f_{1}^{\prime}-f_{2}-3 f_{3}+f_{3}^{\prime}\right) \Omega^{2}-2\left(f_{1}+f_{1}^{\prime}-f_{2}\right) u^{\alpha} R_{\alpha \beta} u^{\beta},
$$

$$
\mathcal{P}=p+\frac{1}{3} f_{1} R-\frac{1}{3}\left(2 f_{1}^{\prime}+f_{3}\right) \Omega^{2}-\frac{1}{3}\left(2 f_{1}^{\prime}+4 f_{1}^{\prime \prime}-f_{2}\right) a^{2}+\frac{2}{3}\left(2 f_{1}^{\prime}-f_{1}\right) u^{\alpha} R_{\alpha \beta} u^{\beta},
$$

$$
\mathcal{Q}_{\mu}=\left(f_{1}^{\prime}+2 f_{3}^{\prime}\right) \epsilon_{\mu \lambda \rho \sigma} a^{\lambda} u^{\rho} \Omega^{\sigma}+\left(2 f_{1}+4 f_{3}\right) \Delta_{\mu}^{\rho} R_{\rho \sigma} u^{\sigma},
$$

$$
\mathcal{T}_{\mu \nu}=\left(4 f_{1}^{\prime}+2 f_{1}^{\prime \prime}-2 f_{2}\right) a_{\langle\mu} a_{\nu\rangle}-\frac{1}{2}\left(f_{1}^{\prime}-4 f_{3}\right) \Omega_{\langle\mu} \Omega_{\nu\rangle}+2 f_{1}^{\prime} u^{\alpha} R_{\alpha\langle\mu \nu\rangle \beta} u^{\beta}-2 f_{1} R_{\langle\mu \nu\rangle} .
$$

$$
\mathcal{N}=n+f_{1, \mu} R+\left(f_{2, \mu}+f_{7}+f_{7}^{\prime}\right) a^{2}+\left(f_{3, \mu}-f_{5}+\frac{1}{2} f_{7}\right) \Omega^{2}-f_{7} u^{\alpha} R_{\alpha \beta} u^{\beta},
$$

$$
\mathcal{J}^{\mu}=-\left(f_{5}+f_{5}^{\prime}\right) \epsilon^{\mu \nu \rho \sigma} u_{\nu} a_{\rho} \Omega_{\sigma}+2 f_{5} \Delta^{\mu \rho} R_{\rho \lambda} u^{\lambda}
$$

$$
f_{n}^{\prime} \equiv T f_{n, T}+\mu f_{n, \mu}, f_{n}^{\prime \prime} \equiv T^{2} f_{n, T, T}+2 \mu T f_{n, T, \mu}+\mu^{2} f_{n, \mu, \mu}
$$

## How to compute the susceptibilities

Kubo formulas is how you connect microscopics (e.g. QCD) to macroscopics (thermodynamics, hydrodynamics)

All seven parity-even susceptibilities are given by 2-point equilibrium functions of $\mathrm{T}^{\mu \mathrm{v}}$ and $\mathrm{J} \mathrm{\mu}$ in flat space.

Can calculate all parity-even susceptibilities on the lattice or in holography.

## Example: free fields

Evaluate the one-loop diagram:


Free massless real scalar:

$$
f_{1}=\frac{T^{2}}{144}(1-6 \xi), \quad f_{2}=0, \quad f_{3}=-\frac{T^{2}}{144} .
$$

Free massless Dirac fermion at $\mu=0$ :

$$
f_{1}=-\frac{T^{2}}{144}, \quad f_{2}=-\frac{T^{2}}{24}, \quad f_{3}=-\frac{T^{2}}{288} .
$$

## Application: hydro with $O(1)$ external magnetic field

$$
\begin{gathered}
\qquad \nabla_{\mu} T \mu v=F v \lambda J_{\lambda} \quad \text { diffeomorphism invariance } \\
\nabla_{\mu} J \mu=0 \quad \text { gauge invariance } \\
T^{\mu v}=T^{\mu v}{ }_{e q}+T^{\mu v} v_{\text {non-eq }}, \quad J \mu=J \mu_{e q}+J \mu_{\text {non-eq }} \\
\text { get from equilibrium } W[g, A]=\int p+O(\partial) \\
\text { e.g. } J \mu_{\text {eq }}=\rho u^{\mu}-\nabla \lambda M \lambda \mu
\end{gathered}
$$

## Application: hydro with $O(1)$ external magnetic field

$$
\begin{gathered}
\nabla_{\mu} T \mu v=F v \lambda J_{\lambda} \\
\nabla_{\mu} J \mu=0 \\
T^{\mu v}=T \mu v e q+T^{\mu v} v_{\text {non-eq }}, \quad J \mu=J \mu_{\text {eq }}+J \mu_{\text {non-eq }}
\end{gathered}
$$

vanish in equilibrium, depend on $\partial_{\mu}, B_{\mu}, E_{\mu}, \eta, \zeta, \ldots$

## Application: hydro with $\mathrm{O}(1)$ external magnetic field

For P-invariant conducting fluid in 3+1dim:

- one thermodynamic susceptibility $\mathrm{M}_{\Omega}=\mathrm{f}_{5}$
- two shear viscosities ( $\perp$ and || to B)
- three bulk viscosities
- two electrical conductivities ( $\perp$ and $|\mid$ to B)
- two Hall viscosities ( $\perp$ and || to B)
- one Hall conductivity

Eleven coefficients total:
1 thermodynamic, non-dissipative
3 non-equilibrium, non-dissipative
7 non-equilibrium, dissipative

## Application: hydro with $\mathrm{O}(1)$ external magnetic field

$$
\begin{aligned}
& \mathcal{E}=-p+T p_{, T}+\mu p_{, \mu}+\left(T M_{\Omega, T}+\mu M_{\Omega, \mu}-2 M_{\Omega}\right) B \cdot \Omega, \\
& \mathcal{P}=p-\frac{4}{3} p_{B^{2}} B^{2}-\frac{1}{3}\left(M_{\Omega}+4 M_{\Omega, B^{2}} B^{2}\right) B \cdot \Omega-\zeta_{1} \nabla \cdot u-\zeta_{2} b^{u} b^{\nu} \nabla_{\mu} u_{\nu}, \\
& \mathcal{Q}^{\mu}=-M_{\Omega} \epsilon^{\mu \nu \rho \sigma} u_{\nu} \partial_{\sigma} B_{\rho}+\left(2 M_{\Omega}-T M_{\Omega, T}-\mu M_{\Omega, \mu}\right) \epsilon^{\mu \nu \rho \sigma} u_{\nu} B_{\rho} \partial_{\sigma} T / T \\
& -M_{\Omega, B^{2}} \epsilon^{\mu \nu \rho \sigma} u_{\nu} B_{\rho} \partial_{\sigma} B^{2}+\left(-2 p_{B^{2}}+M_{\Omega, \mu}-2 M_{\Omega, B^{2}} B \cdot \Omega\right) \epsilon^{\mu \nu \rho \sigma} u_{\nu} E_{\rho} B_{\sigma} \\
& +M_{\Omega} \epsilon^{\mu \nu \rho \sigma} \Omega_{\nu} E_{\rho} u_{\sigma}, \\
& \mathcal{T}^{\mu \nu}=2 p_{, B^{2}}\left(B^{\mu} B^{\nu}-\frac{1}{3} \Delta^{\mu \nu} B^{2}\right)+M_{\Omega, B^{2}} B^{\langle\mu} B^{\nu\rangle} B \cdot \Omega+M_{\Omega} B^{\langle\mu} \Omega^{\nu\rangle} \\
& -\eta_{\perp} \sigma_{\perp}^{\mu \nu}-\eta_{\|}\left(b^{\mu} \Sigma^{\nu}+b^{\nu} \Sigma^{\mu}\right)-b^{\langle\mu} b^{\nu\rangle}\left(\eta_{1} \nabla \cdot u+\eta_{2} b^{\alpha} b^{\beta} \nabla_{\alpha} u_{\beta}\right) \\
& -\tilde{\eta}_{\perp} \tilde{\sigma}_{\perp}^{\mu \nu}-\tilde{\eta}_{\|}\left(b^{\mu} \tilde{\Sigma}^{\nu}+b^{\nu} \tilde{\Sigma}^{\mu}\right), \\
& \mathcal{N}=p_{, \mu}+M_{\Omega, \mu} B \cdot \Omega-m \cdot \Omega, \\
& \mathcal{J}^{\mu}=\epsilon^{\mu \nu \rho \sigma} u_{\nu} \nabla_{\rho} m_{\sigma}+\epsilon^{\mu \nu \rho \sigma} u_{\nu} a_{\rho} m_{\sigma}+\left(\sigma_{\perp} \mathbb{B}^{\mu \nu}+\sigma_{\|} \frac{B^{\mu} B^{\nu}}{B^{2}}\right) V_{\nu}+\tilde{\sigma} \tilde{V}^{\mu} \\
& \text { * In thermodynamic frame, up to } O(\partial) \\
& \Delta^{\mu \nu} \equiv g^{\mu \nu}+u^{\mu} u^{\nu} \quad b^{\mu} \equiv B^{\mu} / B \\
& \sigma^{\mu \nu} \equiv \Delta^{\mu \alpha} \Delta^{\nu \beta}\left(\nabla_{\alpha} u_{\beta}+\nabla_{\beta} u_{\alpha}-\frac{2}{3} \Delta_{\alpha \beta} \nabla \cdot u\right) \\
& \tilde{\sigma}^{\mu \nu} \equiv \frac{1}{2 B}\left(\epsilon^{\mu \lambda \alpha \beta} u_{\lambda} B_{\alpha} \sigma_{\beta}{ }^{\nu}+\epsilon^{\nu \lambda \alpha \beta} u_{\lambda} B_{\alpha} \sigma_{\beta}{ }^{\mu}\right) \\
& \mathbb{B}^{\mu \nu} \equiv \Delta^{\mu \nu}-b^{\mu} b^{\nu} \quad \Sigma^{\mu} \equiv \mathbb{B}^{\mu \lambda} \sigma_{\lambda \rho} b^{\rho} \\
& V^{\mu} \equiv E^{\mu}-T \Delta^{\mu \nu} \partial_{\nu}(\mu / T) \\
& \tilde{v}^{\mu} \equiv \epsilon^{\mu \nu \rho \sigma} u_{\nu} B_{\rho} v_{\sigma} / B \\
& m^{\mu}=\left(2 p_{, B^{2}}+2 M_{\Omega, B^{2}} B \cdot \Omega\right) B^{\mu}+M_{\Omega} \Omega^{\mu}
\end{aligned}
$$

## Application：hydro with $\mathrm{O}(1)$ external magnetic field

Inequality constraints on n＇s，乙＇s，o＇s from 2－nd law
Equality constraints on n＇s，そ＇s，o＇s from Onsager relations
Eigenmodes：collective cyclotron modes，sound，diffusion，．．．
Express n＇s，乙＇s，o＇s in terms of $\left\langle T_{\mu v} T_{\alpha \beta}\right\rangle,\left\langle T_{\mu \nu} J_{a}\right\rangle,\left\langle J_{\mu} J_{a}\right\rangle$
Transport coefficients for P －violating fluids
Hernandez，PK 1703.08757
Huang，Sedrakian，Rischke 1108.0602
Finazzo，Rougemont，Marrochio，Noronha 1412.2968

## Application: Maxwell equations in matter

Equilibrium generating functional $\mathrm{W}\left[\mathrm{g}_{\mu v}, \mathrm{~A}_{\mu}\right]=$ Equilibrium effective action $\mathrm{S}\left[\mathrm{g}_{\mu v}, \mathrm{~A}_{\mu}\right]$

In the vacuum:

$$
S_{\mathrm{eff}}[g, A]=\int d^{d+1} x \sqrt{-g}\left[-\frac{1}{4} F_{\mu \nu}^{2}\right]
$$

$\delta_{A} S_{\text {eff }}=0 \Rightarrow$ Maxwell equations: $J \mu=0$, or $\nabla_{\mathrm{V}} \mathrm{F}^{\mu v}=0$.

## Application: Maxwell equations in matter

Equilibrium generating functional $\mathrm{W}\left[\mathrm{g}_{\mu v}, \mathrm{~A}_{\mu}\right]=$
Equilibrium effective action $\mathrm{S}\left[\mathrm{g}_{\mu v}, \mathrm{~A}_{\mu}\right]$

In matter:

$$
S_{\mathrm{eff}}[g, A]=\int d^{d+1} x \sqrt{-g}\left[-\frac{1}{4} F_{\mu \nu}^{2}+\mathcal{F}_{\mathrm{m}}\left[T, \mu, E^{2}, B^{2}, B \cdot \Omega, \ldots\right]\right]
$$

$\delta_{A} S_{\text {eff }}=0 \Rightarrow$ Maxwell equations: $J^{\mu}=0$, or $\nabla_{V} H^{\mu v}=n u^{\mu}$.

$$
\begin{aligned}
H^{\mu \nu} & \equiv F^{\mu \nu}-M_{m}^{\mu \nu} \\
n & \equiv \partial \mathcal{F}_{\mathrm{m}} / \partial \mu
\end{aligned}
$$

## Application: Maxwell equations in matter

Equilibrium generating functional $\mathrm{W}\left[\mathrm{g}_{\mu v}, \mathrm{~A}_{\mu}\right]=$ Equilibrium effective action $\mathrm{S}\left[\mathrm{g}_{\mu v}, \mathrm{~A}_{\mu}\right]$

Equations to solve:

$$
\begin{aligned}
& \nabla_{\mu} T^{\mu \nu}=F^{\lambda \nu} J_{\mathrm{ext} \lambda} \\
& J^{\mu}+J_{\mathrm{ext}}^{\mu}=0 \\
& \epsilon^{\mu \nu \alpha \beta} \nabla_{\nu} F_{\alpha \beta}=0
\end{aligned}
$$

This is relativistic MHD, with 11 transport coefficients

## MHD vs hydro in external B-field

- MHD has the same 11 transport coef-s (7 are dissipative)
- MHD has the same entropy current
- MHD has the same Kubo formulas for viscosities
- MHD has different Kubo formulas for conductivities

$$
\begin{aligned}
& \frac{1}{\omega} \operatorname{Im} G_{E_{z} E_{z}}^{\mathrm{ret}}(\omega, \mathbf{k}=0)=\rho_{\|} \\
& \frac{1}{\omega} \operatorname{Im} G_{E_{x} E_{x}}^{\mathrm{ret} .}(\omega, \mathbf{k}=0)=\rho_{\perp} \\
& \frac{1}{\omega} \operatorname{Im} G_{E_{x} E_{y}}^{\mathrm{ret}}(\omega, \mathbf{k}=0)=-\tilde{\rho}_{\perp} \operatorname{sign}\left(B_{0}\right)
\end{aligned}
$$

$$
\begin{array}{r}
\sigma_{a b} \equiv \sigma_{\perp} \delta_{a b}+\tilde{\sigma} \epsilon_{a b} \\
\left(\sigma^{-1}\right)_{a b}=\rho_{\perp} \delta_{a b}+\tilde{\rho}_{\perp} \epsilon_{a b} \\
\rho_{\|} \equiv 1 / \sigma_{\|}
\end{array}
$$

- MHD has different eigenmodes (e.g. Alfven waves)


## Questions

There is more to thermodynamics than knowing $p(T, \mu)$. Compute in lattice QCD \& in AdS?

Well-posedness of MHD a la Israel-Stewart?

Transport coef-s in B-field at weak vs strong coupling? Physical implications?

Statistical fluctuations, aggravated by the B-field?
There is a "dual" formulation of MHD in terms of the magnetic flux. Relation to "conventional" MHD underexplored.

Thank you!

